# Preservation of Evidence of Ancient Environments and Life on Mars



Mars today



Early Mars ?

Dave Des Marais NASA Ames Research Center Key factors affecting biosignature preservation & abundance on Earth

**Paleo-productivity** 

Transport and burial

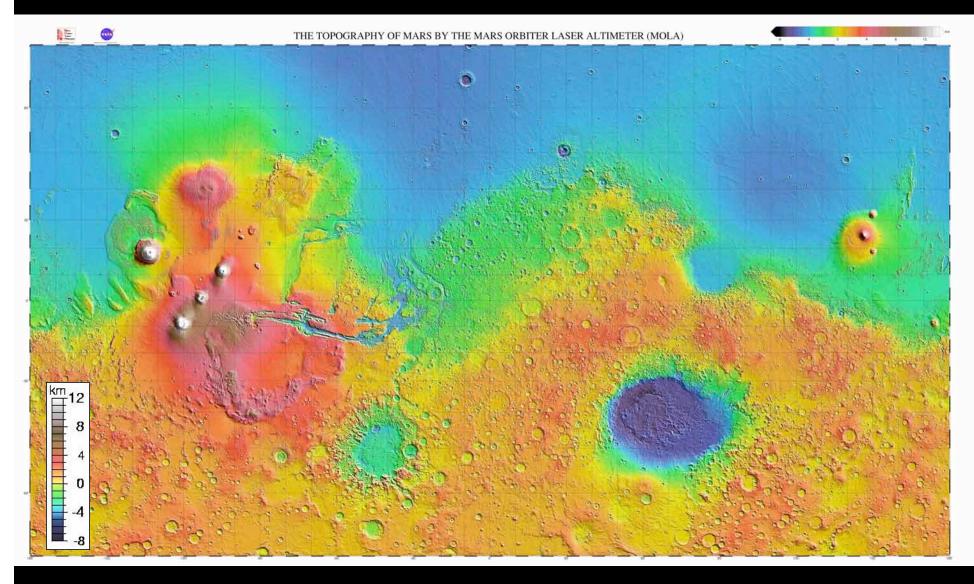
Sedimentary redox conditions

Mineralogy

Lithification

Subsequent alteration / destruction

## Mars Global Surveyor MOLA Topography



### Martian attributes relevant to perservation

Noachian: Active tectonics, magnetic field, nearsurface water was relatively abundant and persistent (also a northern lowlands large water body?)

Lower elevations and deeper subsurface environments enchance persistence of aqueous conditions (thus favor habitable environments, sediment lithification)

Phyllosilicate deposits in reducing environments enhance organic abundances and preservation

Implications for relative preservation potential of the four "finalist" MSL sites



### Preservation of Evidence about Ancient Mars

Noachian-Hesperian environments and processes

Persistence of ancient aqueous environments

Noachian deposits in aqueous environments

Phyllosilicate-rich deposits and organic matter contents

Preservation potential of the four "finalist" MSL sites

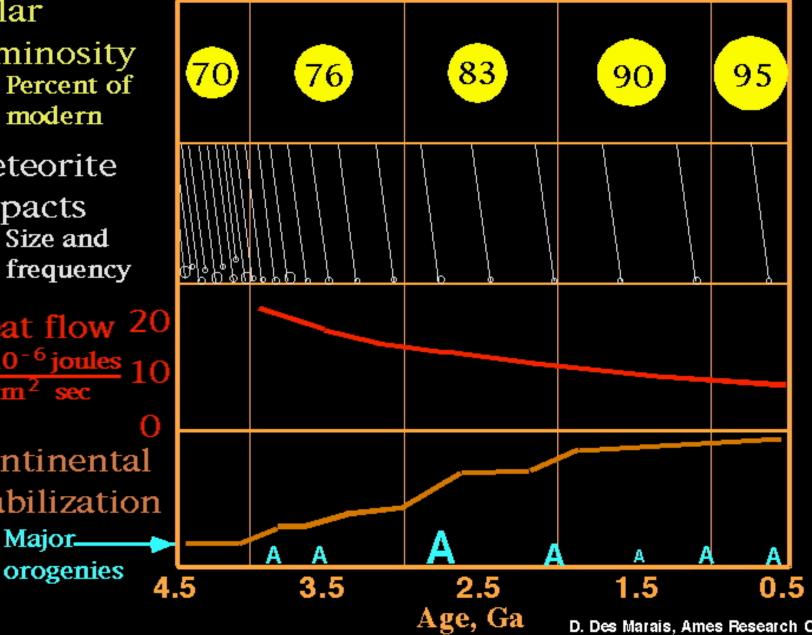
#### Evolution of Earth's Early Environment

Solar luminosity Percent of modern

Meteorite impacts Size and frequency

Heat flow 20 <sup>6</sup> joules 10 cm<sup>2</sup> sec

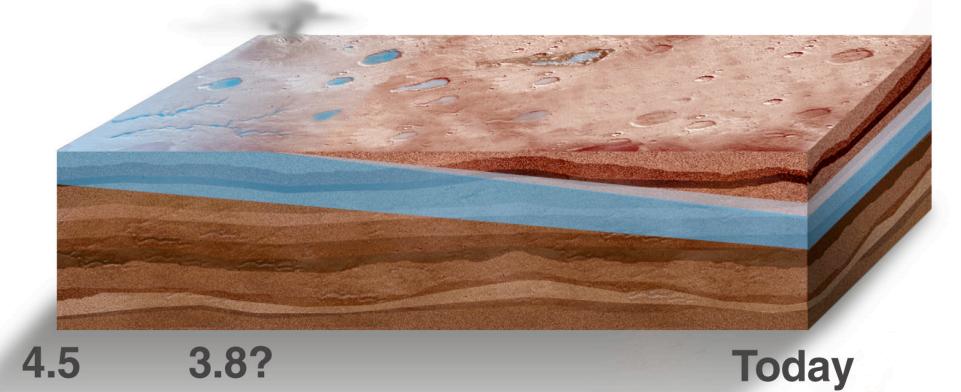
Continental stabilization Major\_



D. Des Marais, Ames Research Center

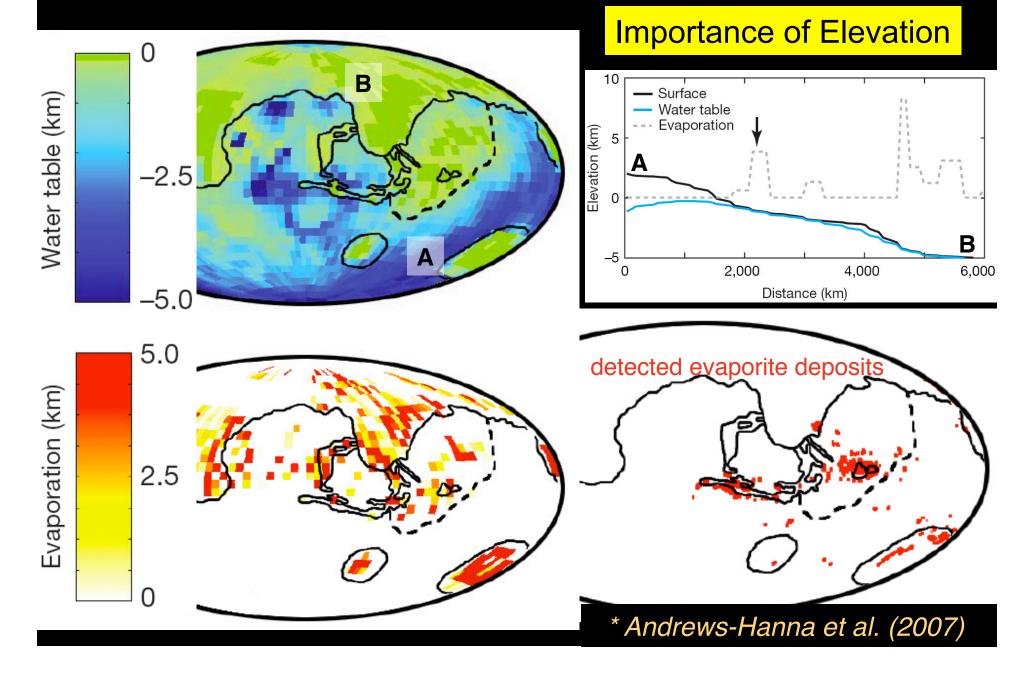
## Conditions That Could Sustain Life on Mars: Changes Over the Eons

Importance of Age

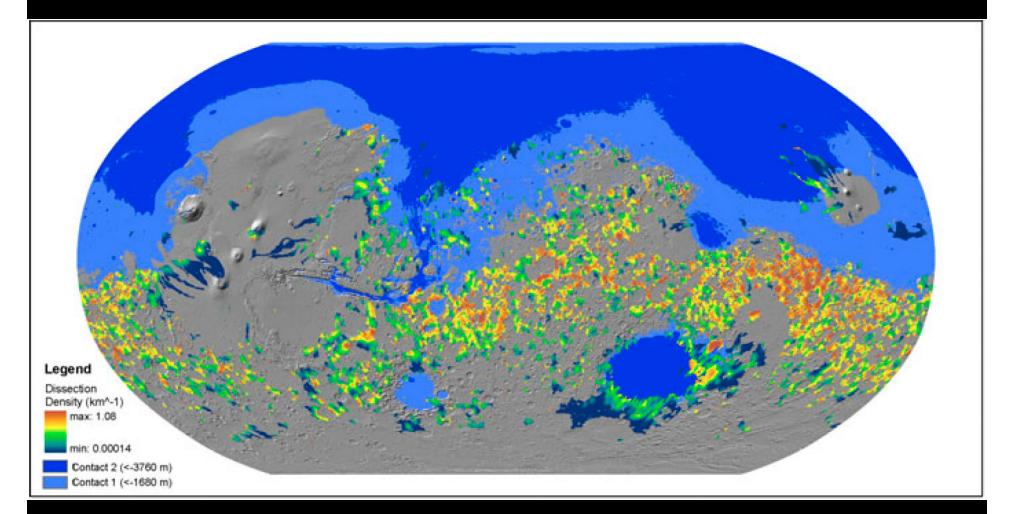


**Billions of Years Ago** 

### Meridiani Planum and the Global Hydrology of Mars \*

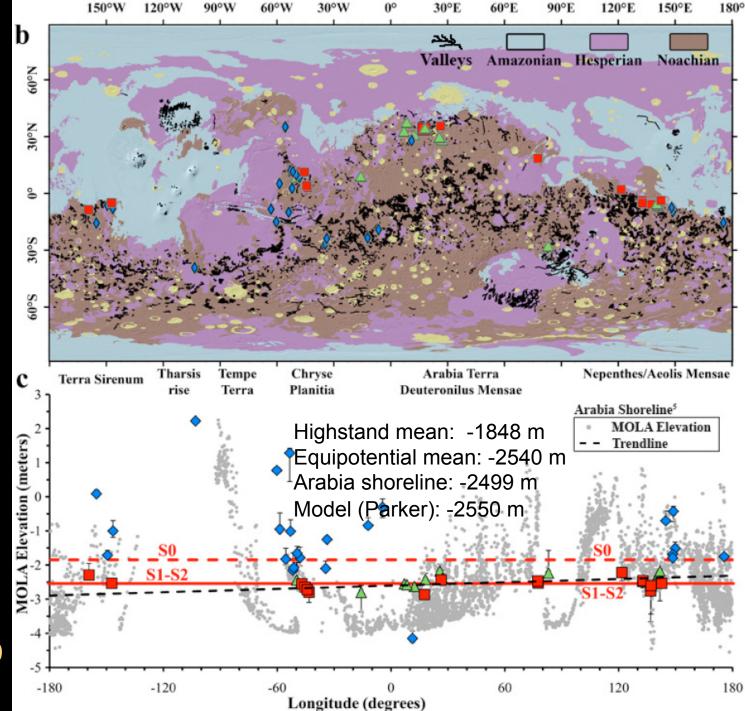


## Stream Systems



#### Stepinski & Luo, LPSC41, 2010

## Deltas



Di Achille & Hynek, LPSC41, 2010 Apollinaris Patera (volcano)

4

Northern Lowlands

\*

Southern Highlands

Gusev Crater ~180 km diameter Ma'Adim Vallis

X

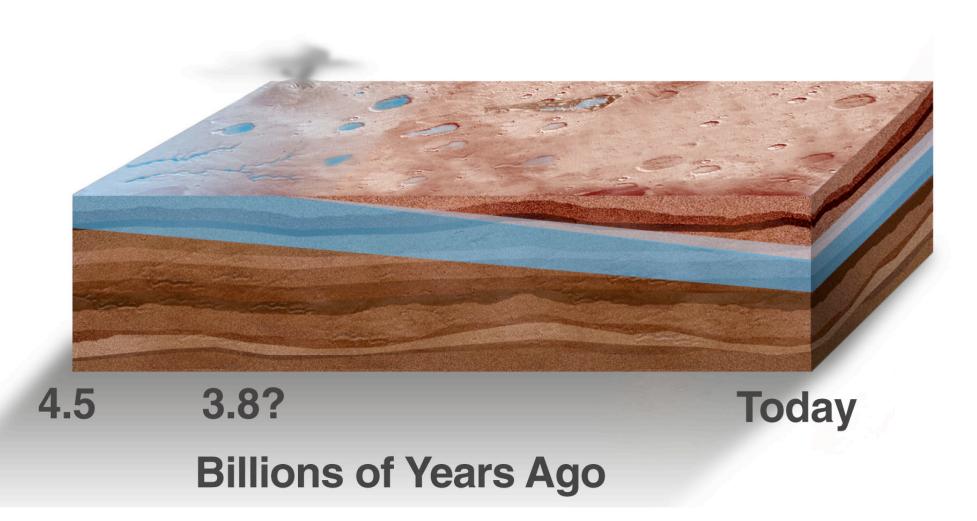
▲Orbiter view (MRO HiRISE)

Husband Hill and Inner Basin

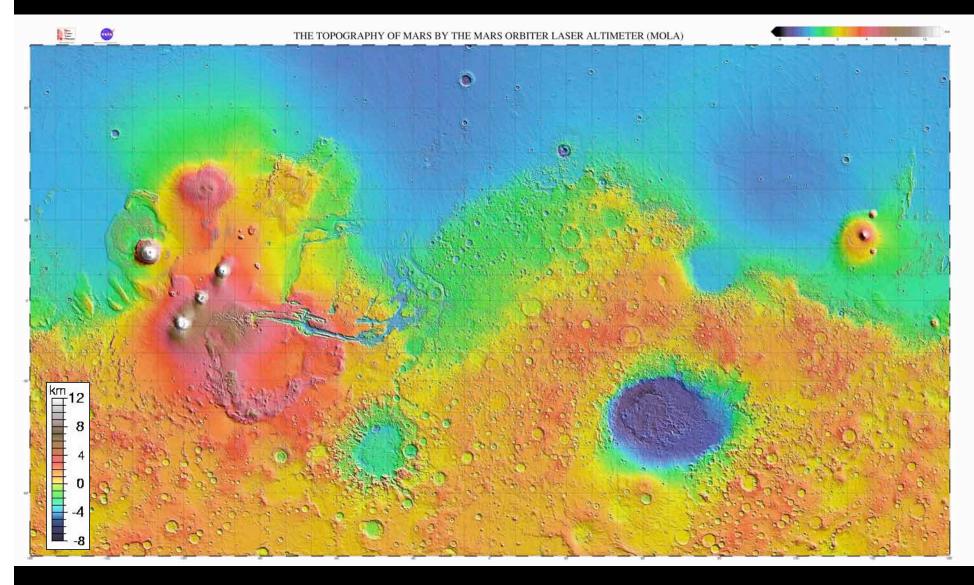
ground view (MER Pancam) Vesicular basalts: water-rich magma Explosive volcanism: volatile-rich Bomb sag: bomb impacting wet sediment Ferric sulfate-rich deposits: hydrothermal/fumarolic/acidic Pure silica: sinter/acid leaching Carbonate rich bedrock

## Conditions That Could Sustain Life on Mars: Changes Over the Eons

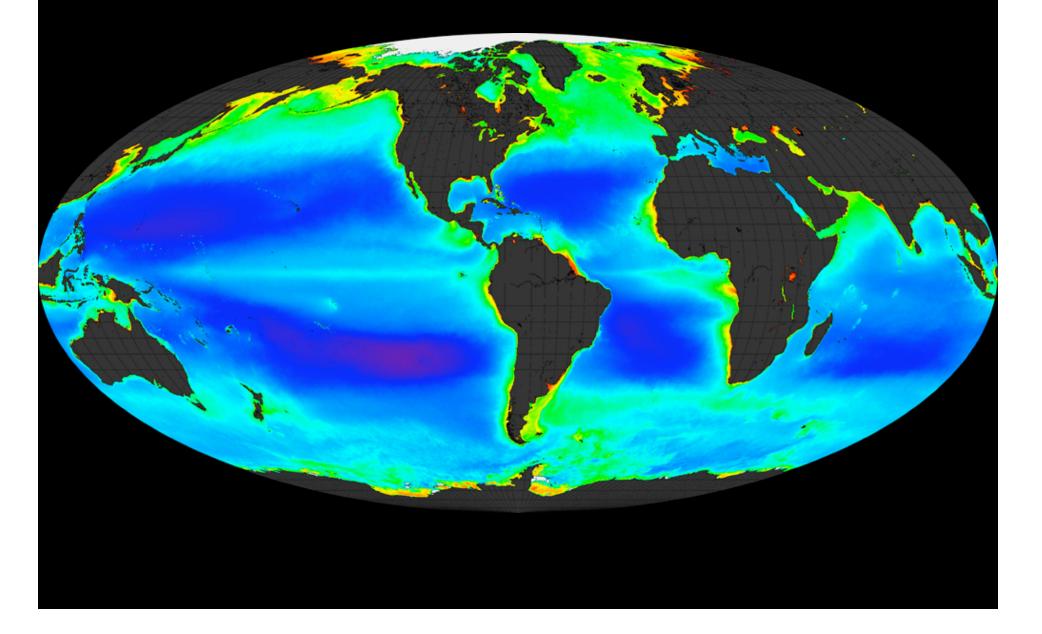
#### Importance of Age and Depth in Crust



## Mars Global Surveyor MOLA Topography



### Marine Chlorophyll Abundances (Low Moderate High)

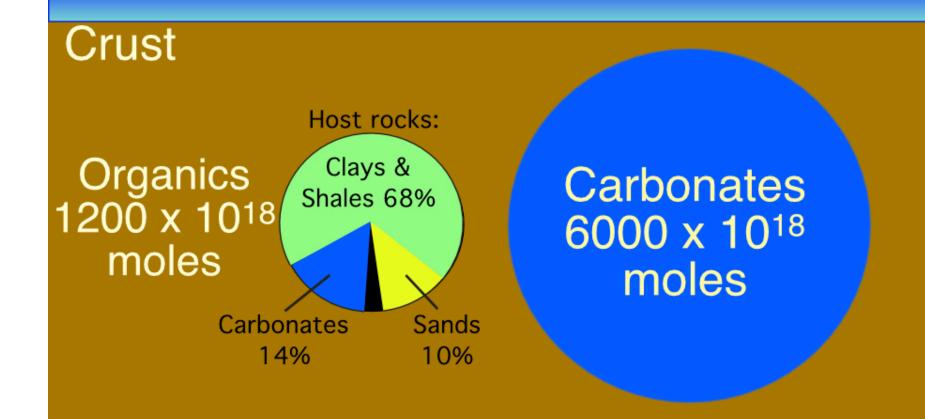


## Earth's Carbon Budget

Biosphere, Oceans and Atmosphere

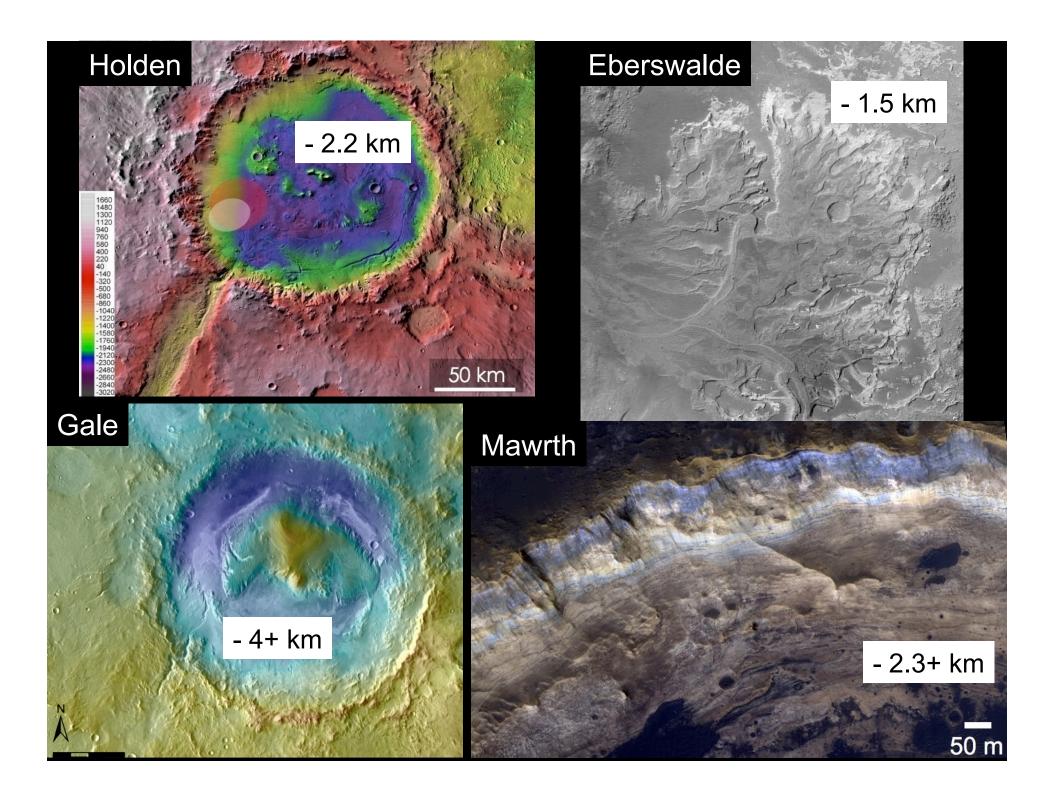
Mantle

• 3.7 x 10<sup>18</sup> moles



~20,000 x 1018 moles

Late Proterozoic Shale	TOC, mg/g			
Wyman	0.9	to	2.3	_
Chapel Is.	< 0.02	to	0.45	
Cijara	1.2			Key Factors:
Fuentes	0.8			Paleo-productivity
Klatyspyt	207			
Pusa	1.6			Transport & burial
Timperley	0.3			Sedimentary redox
Sheepbed	1.4	to	10	Mineralogy
Luoquan	0.3	to	3.9	Lithification
Twitya	0.1	to	2.4	Later alteration
Tapley Hill	2.6			
Woocalla	2.6			
River Wakefield	0.3			
Visingso	0.4	to	2.0	
Little Dal	0.4	to	6.9	
Sayunei	< 0.02	to	0.06	
Chuar	0.2	to	29.6	
Red Pine	2.2			
Kuktur	7.9			



**Preservation of Evidence about Ancient Mars** Noachian-Hesperian environments and processes Persistence of ancient aqueous environments Noachian deposits in aqueous environments Phyllosilicate-rich deposits and organic matter contents Preservation potential of the four "finalist" MSL sites Age of deposits: mid-Noachian to early Hesperian Water persistence: effects of elevation and geologic age Allochthonous vs autochthonous phyllosilicates (Kennedy work, redox state, nature of organics [e.g., plant lignin can survive transport; can microbial components?]) Redox state of deposit Lithification: rate and extent